



Standard Test Method for Calorimetric Determination of Hemispherical Emittance and the Ratio of Solar Absorptance to Hemispherical Emittance Using Solar Simulation¹

This standard is issued under the fixed designation E434; the number immediately following the designation indicates the year of original adoption or, in the case of revision, the year of last revision. A number in parentheses indicates the year of last reapproval. A superscript epsilon (ϵ) indicates an editorial change since the last revision or reapproval.

1. Scope

1.1 This test method covers measurement techniques for calorimetrically determining the ratio of solar absorptance to hemispherical emittance using a steady-state method, and for calorimetrically determining the total hemispherical emittance using a transient technique.

1.2 *This standard does not purport to address all of the safety concerns, if any, associated with its use. It is the responsibility of the user of this standard to establish appropriate safety and health practices and determine the applicability of regulatory limitations prior to use.*

2. Referenced Documents

2.1 *ASTM Standards:*²

E349 [Terminology Relating to Space Simulation](#)

3. Summary of Test Method

3.1 In calorimetric measurements of the radiative properties of materials, the specimen under evaluation is placed in a vacuum environment under simulated solar radiation with cold surroundings. By observation of the thermal behavior of the specimen the thermophysical properties may be determined by an equation that relates heat balance considerations to measurable test parameters.

3.2 In a typical measurement, to determine α/ϵ as defined in Definitions E349, the side of the specimen in question is exposed to a simulated solar source, through a port having suitable transmittance over the solar spectrum. This port, or window, must be of sufficient diameter that the specimen and radiation monitor will be fully irradiated and must be of sufficient thickness that it will maintain its strength without deformation under vacuum conditions. The radiant energy absorbed by the specimen from the solar source and emitted by

the specimen to the surroundings cause the specimen to reach an equilibrium temperature that is dependent upon the α/ϵ ratio of its surface.

3.3 In the dynamic radiative method of measuring total hemispherical emittance, the specimen is heated with a solar simulation source and then allowed to cool by radiation to an evacuated space chamber with an inside effective emittance of unity. From a knowledge of the specific heat of the specimen as a function of temperature, the area of the test specimen, its mass, its cooling rate, and the temperature of the walls, its total hemispherical emittance may be calculated as a function of temperature.

4. Apparatus

4.1 The main elements of the apparatus include a vacuum system, a cold shroud within the vacuum chamber, instrumentation for temperature measurement, and a solar simulator.

4.2 The area of the thermal shroud shall not be less than 100 times the specimen area (controlled by the specimen size). The inner surfaces of the chamber shall have a high solar absorptance (not less than 0.96) and a total hemispherical emittance of at least 0.88 (painted with a suitable black paint),³ and shall be diffuse. Suitable insulated standoffs shall be provided for suspending the specimen. Thermocouple wires shall be connected to a vacuumtight fitting where the temperature of feedthrough is uniform. Outside of the chamber, all thermocouples shall connect with a fixed cold junction.

4.3 The chamber shall be evacuated to a pressure of 1×10^{-6} torr (0.1 mPa) or less at all times.

4.4 The walls of the inner shroud shall be in contact with coolant so that their temperature can be maintained uniform at all times.

4.5 A shutter shall be provided in one end of the chamber which can be opened to admit a beam of radiant energy from a solar simulator. When open, this shutter shall provide an aperture admitting the full simulator beam. When the shutter is

¹ This test method is under the jurisdiction of ASTM Committee E21 on Space Simulation and Applications of Space Technology and is the direct responsibility of Subcommittee E21.04 on Space Simulation Test Methods.

Current edition approved Oct. 1, 2015. Published November 2015. Originally approved in 1971. Last previous edition approved in 2010 as E434 – 10. DOI: 10.1520/E0434-10R15.

² *Annual Book of ASTM Standards*, Vol 15.03.

³ Nextel Brand Velvet Coating 401-C10 Black, available from Reflective Products Div., 3M Co., has been found to be satisfactory.

closed, all rays emitted by the specimen shall be intercepted by a blackened surface at the coolant temperature (the shutter must be at least conductively coupled to the shroud).

4.6 The vacuum chamber shall be provided with a fused silica window large enough to admit the simulator beam and uniformly irradiate the entire specimen projected area. This window shall have high transmittance through the solar spectrum wavelength region. The chamber shall be provided with a vacuumtight sleeve for opening and closing the shutter and standard vacuum fittings for gaging, bleeding, leak testing, and pumping. If low α/ε specimens are to be measured, the solid angle subtended by the port from the specimen should be small (dependent upon desired accuracy). If flat specular specimens are to be measured, the port plane should be canted with respect to the specimen plane to eliminate multiple reflections of the simulator beam. Multiple reflections could result in as much as a 7 % apparent increase in α/ε .

4.7 The solar simulator should duplicate the extraterrestrial solar spectrum as closely as possible. A beam irradiance of at least 7000 W/m^2 at the specimen plane shall be available from the solar simulator (~ 5 solar constants). This irradiance may be required to raise the temperature of certain specimens to a desired level.

5. Coating Requirements

5.1 Any type of coating may be tested by this test method provided its structure remains stable in vacuum over the temperature range of interest.

5.2 For high emittance specimens the accuracy of the measurements is increased if only one surface of the substrate is coated with the specimen coating in question. The remaining area of the substrate shall be coated with a low emittance material of known hemispherical emittance (such as evaporated aluminum or evaporated gold).

5.3 The thickness and density of the coating shall be measured and its heat capacity calculated from existing references (see Refs (1) and (2)).⁴

6. Specimen Preparation

6.1 The substrates used for the measurements described here shall be of a material whose specific heat as a function of temperature can be found in standard references (for example, OFHC copper or a common aluminum alloy such as 6061-T6) (Ref (1)).

6.2 The substrate shall be machined from flat stock and to a size proportioned to the working area of the chamber.

6.3 Each specimen shall be drilled with a set of holes, near the edge, through which suspension strings are to be inserted.

6.4 Each substrate shall be drilled with two small shallow holes in the back for thermocouples.

6.5 Ideally the back and sides of the substrate shall be buffed and polished and one uninsulated thermocouple inserted

in the back of the specimen (one wire in each hole). One of these wires shall be peened into each hole.

6.6 A low-emittance coating shall be applied to the back and sides of the substrate and to the thermocouple wires for several inches at the specimen end.

6.7 The substrates shall be coated with the material in question. The coating shall be of sufficient thickness so as to be opaque. (This will avoid any substrate effects.)

6.8 The specimens shall be suspended from the top of the shroud by means of thread or string. These strings shall be of small diameter, low thermal conductivity, and low emittance in order to minimize heat losses through the leads.

6.9 An alternative method of specimen mounting (mass dependent) shall be to suspend the specimens by their own small wire thermocouple leads. In this case the thermocouple holes shall be drilled as before but radially around the edge. The suspension holes may also be eliminated in this case.

7. Procedure

7.1 Suspend the test specimen in the chamber normal to the incident solar radiation, but geometrically removed from the central axis of the chamber so that radiation from the specimen to the chamber walls is not specularly reflected back to the specimen. Since the chamber walls are designed to be cold and highly absorbing, first reflections from the walls are usually all that need be considered.

7.2 Determine the simulated solar irradiance incident on the specimen with a suitable radiometric device such as a commercial thermopile radiometer or a black monitor sample of known α/ε which may be suspended similarly to the test specimen within the incident beam of simulated solar radiation. Take care in the latter case that the irradiance and spectral distribution of the incident energy is the same for both specimen and monitor.

7.3 Then close the system and start the evacuation and cooling of the shroud (see Ref (3) for a typical system). Maintain a pressure of 1×10^{-6} torr (0.1 mPa) or less and the walls of the chamber must be at coolant temperature. Record the specimen, monitor, and shroud temperatures.

7.4 When the specimen has reached thermal equilibrium, that is, when the specimen temperature becomes constant with constant surrounding conditions, shut off the solar simulator. When specimens of large thermal mass are used, carefully evaluate the $\Delta T/\Delta t = 0$ conditions, that is, the Δt chosen should be dependent on the specimen time constant.

7.5 Close the moveable door in the shroud and allow the specimens to cool to a desired temperature. Measure the specimen temperature as a function of time and calculate the rates of change of the temperature.

8. Calculation

8.1 Calculate the $\alpha_{es}/\varepsilon (T_1)$ ratio from the following equation:

$$\frac{\alpha_{es}}{\varepsilon (T_1)} = \frac{A_r}{A_p E} \sigma \left(T_1^4 - \frac{\varepsilon (T_0)}{\varepsilon (T_1)} T_0^4 \right) \quad (1)$$

⁴ The boldface numbers in parentheses refer to the list of references appended to this method.

where:

- α_{es} = Effective solar absorptance relative to the illuminating source,
- $\varepsilon(T_0)$ = hemispherical emittance of the specimen at Temperature T_0 ,
- $\varepsilon(T_1)$ = hemispherical emittance of the specimen at Temperature T_1 ,
- σ = Stefan-Boltzmann constant,
- A_p = projected area of the specimen exposed to solar radiation,
- E = incident total irradiance,
- T_1 = specimen equilibrium temperature with simulated solar radiation,
- T_0 = chamber wall temperature with solar source off, and
- A_T = total radiating area of the specimen.

8.2 This equation is derived in the following manner: If a specimen coated on all sides with the material in question, with a projected area as viewed in the direction of irradiation, A_p , a total area, A_T , effective simulated solar absorptance, α_{es} , emittance at T_1 , $\varepsilon(T_1)$, and specific heat c_p is suspended in an evacuated high absorptance isothermal cold-walled chamber and exposed to a simulated solar irradiance, E , the rate of temperature change can be determined by evaluating the heat balance equation. The energy balance of an irradiated specimen emitting radiant energy in a vacuum is given by the following equation (assuming parasitic heat losses can be ignored):

$$mc_p \left(\frac{dT}{dt} \right) = A_p \sigma E_p + E_p - A_i \varepsilon(T_1) \sigma T_1^4 + A_r \alpha_{ir}(T_0) \sigma T_0^4 \quad (2)$$

where $E_p = A \varepsilon \sigma T_2^4$, the thermal radiation from the port. To determine the incident thermal radiation, E_p , see Ref (3). The last term, $A_r \alpha_{ir}(T_0) \sigma T_0^4$, is the amount of heat energy absorbed by the sample from the chamber walls. Kirchoff's law tells us that at a given temperature the infrared absorptance is equal the infrared emittance. This means that it will emit as much heat as it absorbs from a black body at the same temperature as the sample. Therefore, to know how much energy is absorbed by the sample from the shroud walls we must know the infrared absorptance (and hence the emittance) of the sample at the temperature of the shroud wall. The infrared absorptance at T_0 $\alpha_{ir}(T_0)$, by Kirchoff's law is equal to the infrared emittance of the sample at that temperature so we can write:

$$mc_p \left(\frac{dT}{dt} \right) = A_p \sigma E_p + E_p - A_i \varepsilon(T_1) \sigma T_1^4 + A_r \varepsilon(T_0) \sigma T_0^4 \quad (3)$$

If E_p is eliminated from Eq 2 when an equilibrium temperature is reached, $mc_p(dT/dt) = 0$, and, From Eq 2, solving for the α/ε ratio we obtain

$$\frac{\alpha_{es}}{\varepsilon(T_1)} = \frac{A_r}{A_p E} \sigma \left(T_1^4 - \frac{\varepsilon(T_0)}{\varepsilon(T_1)} T_0^4 \right) \quad (4)$$

Eq 4 is used to calculate the $\alpha_{cs}/\varepsilon(T_1)$ ratio when the parameters A_T , E , and A_p are determined and the equilibrium temperature is measured.

8.3 If the source is blocked by the shutter and the specimen loses energy only by radiation, the energy balance equation becomes:

$$mc \left(\frac{dT}{dt} \right) = A_i \varepsilon(T_1) \sigma T_1^4 - A_r \alpha_{ira}(T_0) \sigma T_0^4 + Q_{ll} + Q_{rg} - Q_{ts} \quad (5)$$

Where Q_{ll} and Q_{rg} represent the heat losses from the support leads and the heat lost from the residual gasses in the vacuum chamber, respectively. The last term Q_{ts} is any heat input from the temperature sensor. See Ref (4) and Ref (5) for a treatment of the lead loss and residual gas heat loss terms.

8.4 If the term T_0^4 is neglected, and the parasitic heat losses and gains can be ignored, the above equation can be integrated and expanded into:

$$\varepsilon(T_1) = \frac{(m_s c_s + m_c c_c)}{3 \sigma A_i \Delta t} \left(\frac{1}{T_1^3} - \frac{1}{T_2^3} \right) \quad (6)$$

where:

- m_s = mass of the substrate,
- m_c = mass of the coating,
- c_s = thermal capacitance of the substrate,
- c_c = thermal capacitance of the coating,
- T = temperature of the specimen, and
- Δt = change in time from T_1 to T_2 and magnitude such that c_s and c_c may be assumed constant over small temperature ranges.

When the temperature decay is recorded with time, then the total hemispherical emittance of the sample can be determined with Eq 5 or Eq 6. The use of Eq 6 is preferable since Eq 5 involves the experimental determination of two quantities (dT/dt and T^4), thereby introducing more possible errors than in Eq 6.

8.5 Data from specimens which are coated on one side only shall be reduced by use of the following equation:

$$\varepsilon(T)_c = \frac{(m_s c_s + m_c c_c)}{3 \sigma A_c \Delta t} \left(\frac{1}{T_1^3} - \frac{1}{T_2^3} \right) - \frac{\varepsilon_s (A_T - A_c)}{A_c} \quad (7)$$

where:

- ε_s = total hemispherical emittance of substrate,
- A_c = area of coating, and
- ε_c = total hemispherical emittance of coating.

8.6 To obtain an α/ε measurement or an effective solar absorptance, α , for a specimen coated only on one side, one must consider the following expression:

$$A_T \varepsilon_T = A_c \varepsilon_c + A_s \varepsilon_s \quad (8)$$

where:

- A_T, A_c, A_s = total area, area of the coating, and uncoated area of the substrate, respectively, and
- $\varepsilon_T, \varepsilon_c, \varepsilon_s$ = total hemispherical emittance of the specimen, coating, and substrate respectively.

Rearrangement shows that:

$$\varepsilon_T = (A_c \varepsilon_c + A_s \varepsilon_s) / A_T \quad (9)$$

Multiplying the α/ε value obtained from Eq 4 by ε_T (at the same temperature of equilibrium) obtained from Eq 9 will give the solar absorptance, α . In order to acquire the (α/ε) coating, divide the α_s value by ε_c (already measured in a transient cool down).

9. Report

9.1 The report should include the methods used for temperature and irradiance measurements, and the actual data used for the calculations.

9.2 A complete characterization of the specimen shall be given whenever possible. This shall include specimen dimensions, specimen composition, coating thickness and composition, surface roughness, and surface contamination, and any other conditions which may be considered pertinent.

9.3 In an α/ε type of measurement, the total exposure time and level of irradiance, and spectral distribution of the incident flux shall also be reported.

10. Uncertainty Analysis

10.1 Many potential errors exist in the calorimetric determination of radiative properties. If it is assumed that the major uncertainties encountered in these calorimetric measurements are systematic rather than random, they will add in a linear manner and the total uncertainty can be expressed as:

$$\delta\varepsilon_s/\varepsilon_s = (\delta\varepsilon_s/\varepsilon_s)_{conv} + (\delta\varepsilon_s/\varepsilon_s)_q + (\delta\varepsilon_s/\varepsilon_s)_R + (\delta\varepsilon_s/\varepsilon_s)_{HL} \quad (10)$$

for emittance, and as

$$\frac{\delta\alpha/\varepsilon}{\alpha/\varepsilon} = \left(\frac{\delta\alpha/\varepsilon}{\alpha/\varepsilon}\right)_{conv} + \left(\frac{\delta\alpha/\varepsilon}{\alpha/\varepsilon}\right)_R + \left(\frac{\delta\alpha/\varepsilon}{\alpha/\varepsilon}\right)_{HL} + \left(\frac{\delta\alpha/\varepsilon}{\alpha/\varepsilon}\right)_s \quad (11)$$

for the ratio of solar absorptance to hemispherical emittance. The terms on the right of the emittance uncertainty equation can be defined as *conv* the conventional error, *q* the heat measurement error, *R* the extraneous radiation error, and *HL* the heat loss error, respectively. In the uncertainty equation for α/ε the last term, *s*, is defined as the error due to solar simulation where all fractional errors have been previously defined. These uncertainties are discussed in the following paragraphs.

10.2 *Conventional Error*—The conventional error contribution to the total uncertainty involves errors in the measurement of the basic physical quantities of a sample such as the area of the sample, the temperature of the sample, and the enclosure temperature. This error can be expressed as:

$$\left(\frac{\delta\varepsilon/\varepsilon}{\alpha/\varepsilon}\right)_{conv} = \left(\frac{\delta\varepsilon_s}{\varepsilon_s}\right)_{conv} = \frac{\delta A_T}{A_T} \quad (12)$$

$$+ 4 \left(\frac{T_1^4}{T_1^4 - T_0^4}\right) \frac{\delta T_1}{T_1} + 4 \left(\frac{T_0^4}{T_1^4 - T_0^4}\right) \frac{\delta T_0}{T_0}$$

In a given system, each of the quantities in Eq 10 is subject to a varying degree of accuracy; however, the most significant uncertainty occurs in measurement of specimen temperature. The magnitude of the uncertainties of specimen and enclosure temperatures can be determined when thermocouples with known calibrations are utilized. Thermocouples used in these investigations should have a maximum deviation of ± 0.5 K and should read on the 1968 International temperature scale. From Eq 10 a total maximum uncertainty of calculated emittance values is about 2.0 %.

10.3 *Heat Measurement Error*—The error in determination of heat radiation by the test specimen in the enclosure walls

includes the uncertainties in the measurement of Δt , specimen mass, and specific heat of the test specimen which can be expressed as follows:

$$(\delta\varepsilon_s/\varepsilon_s)_q = (\delta m/m) + (\delta c_p/c_p) + (\delta\Delta t/\Delta t) \quad (13)$$

10.3.1 To determine the uncertainty of specimen mass, it is necessary to know the balance manufacturers stated uncertainty. A precision balance from any reputable manufacturer will yield a purely negligible error (on the order of ± 0.003 % for any metallic substrate). The uncertainties of specific heat values as published in the technical literature are dependent upon the material of the substrate; stated values of specific heat are known to a higher degree of accuracy for elemental materials than for most alloys. The published values of c_p for the substrates suggested for use in these investigations are known to within ± 1 %. These uncertainties in c_p can be considered as maximum and result in like uncertainties in calculated values of emittance. For emittance measurements of coatings having an unknown c_p , an elemental substrate should be used. It is evident that the thermal mass of the coating shall not be a significant percentage of the thermal mass of the substrate (less than 2 %). This procedure will minimize the error due to the uncertainty of the thermal capacitance of the coating. For the method described here, the uncertainty in the measurement of time intervals, Δt , is on the order of ± 0.3 s. The maximum uncertainty involved would occur for high ε materials (greater than 0.93). This results in a maximum uncertainty of approximately 3.5 % for calculated values of emittance. (This occurs at relatively high temperatures.)

10.4 *Extraneous Radiation*—Extraneous radiation depends to a great extent on the chamber geometry and the position of the specimen within the chamber. Several possible sources of extraneous radiation that could exist in these measurements are:

- (1) Thermal radiation from the chamber walls,
- (2) Thermal radiation from the port and through the port from the ambient environment,
- (3) Thermal radiation from the specimen reflected back to the specimen from the chamber walls,
- (4) Solar source radiation reflected by the chamber walls onto the specimen,
- (5) Solar source radiation reflected from the specimen to the chamber walls or port and back to the specimen, and
- (6) Radiation from pumps and other heat sources internal to the system which may be viewed by the specimen.

The design of an apparatus for measurement of radiative properties by the calorimetric method must minimize the errors associated with reflection by the enclosure walls of the specimen-emitted energy and by the radiation of energy from the enclosure walls. This minimization of errors is accomplished by making the ratio of the wall area to the specimen area as large as practical, by coating the enclosure walls with a diffusely reflecting highly absorbing surface, and by maintaining the enclosure walls at relatively cold temperature.

10.4.1 Radiation from the ambient environment cannot be easily eliminated, but it is recommended that an evaluation of radiation from and through the porthole be made by measuring

the equilibrium temperature of a specimen with the port open and no solar simulation.

10.4.2 Errors due to (3), (4) and (5) will be minimized by placing the specimen off the geometrical axis of the chamber so that the specular reflections from the surroundings do not fall upon the specimen.

10.5 *Heat Losses Error*—The fractional error due to conductive heat losses (Ref (6)) (thermal conductions through lead wires and thermal conduction through the residual gas in the vacuum chamber) can be expressed as:

$$\left(\frac{\delta a/\varepsilon}{a/\varepsilon}\right)_{HL} = \left(\frac{\delta \varepsilon_s}{\varepsilon_s}\right)_{HL} \leq \frac{\pi N}{A_T \varepsilon} \left(\frac{K \varepsilon_w D^3}{10 \sigma T^3}\right)^{\frac{1}{2}} + \left(\frac{3}{2}\right) \frac{k(T_1 - T_0)v}{\varepsilon \sigma (T_1^4 - T_0^4)} \quad (14)$$

where K is thermal conductivity of N wires, of emittance, ε_w , and diameter, D , k is the Boltzmann constant, v is the number of molecules that impinge on the unit area of the specimen in unit time. The first term shows the contribution from heat loss through the lead wires due to thermal conduction. The differential equation describing the heat flow through semi-infinitely long wires with isothermal cross sections can be expressed as:

$$K \frac{d^2 \theta}{dx^2} \pi r_w^2 = \varepsilon_w \sigma (\theta^4 - T_0^4) 2\pi r_w \quad (15)$$

where:

θ = temperature along wire,

r_w = radius of wire, and

x = distance along wire measured from the specimen.

Upon integration we proceed to:

$$q = \pi N \left[K \sigma \varepsilon_w D^3 \frac{(T_1^5 - 5T_0^4 T_1 + 4T_0^5)}{10} \right]^{\frac{1}{2}} \quad (16)$$

For a specimen possessing a very low emittance at about 200 K, the worst case, the fractional error is 0.4 %. The fractional error due to thermal conduction through the gas in the vacuum chamber is negligible at pressures below 10^{-6} torr (0.1 mPa), even with emittances as low as 0.01, for temperatures above 150 K. (See Ref (7) for temperatures below 150 K.)

10.6 *Solar Simulation Error*:

10.6.1 A fractional error in a/ε measurements due to solar simulation can exist unless care is taken to avoid it. Contributions due to uncertainties in irradiance measurements, uniformity of irradiance, source stability, and spectral mismatch comprise any total error involved.

10.6.2 Typical methods of obtaining a measurement of the simulated solar irradiance is to utilize a thermopile or black monitor sample. Most commercial thermopiles have reported accuracies on the order of ± 2 %. Black monitor specimens have demonstrated good control as long as their radiative

properties are not altered by repeated vacuum cycles, oil backstreaming, or ultraviolet degradation from the light source itself.

10.6.3 Total irradiance measurements should be made at the specimen position, correcting for the transmittance of the port. Spectral measurements shall also be made through the port.

10.6.4 When actual measurements are in progress, care shall be taken to monitor any changes in port transmittance due to deposition of residual material (that is, outgassing material or titanium from sublimation pumps) on the port surface which may make a considerable change in irradiance at the specimen position.

10.6.5 Electrode feed in carbon arc sources, power supply stability, and lamp life in discharge lamps may contribute to instability in energy sources. The recommended procedure is not only to execute good controls on the sources but also to use continuous monitoring instrumentation to document unstable conditions. When the thermal mass of the specimen is large, very short fluctuations in the light source will have little effect upon the results. Any solar simulation system used will not match the solar spectral irradiance perfectly. This spectral mismatch results in an absorptance, α , that is not a true solar absorptance. This absorptance is more correctly called an effective absorptance and depends upon the spectral irradiance of the source. Examples of the magnitude of the errors involved in the calculated effective absorptance, α , when the difference in source spectral irradiance is not considered are given in Table 1 (see Ref (8)). These data were obtained by multiplying the spectral absorptance of the specimen in question incrementally by the spectral irradiance of the source used and integrating the resulting curve to obtain the total effective specimen absorptance. This is the effective absorptance and can be expressed mathematically by the following equation:

$$\alpha_{\text{eff}} = \int_{\lambda_1}^{\lambda_2} E_{\lambda} \alpha(\lambda) d\lambda / \int_{\lambda_1}^{\lambda_2} E_{\lambda} d\lambda \quad (17)$$

where:

E = spectral irradiance,

$\alpha(\lambda)$ = spectral absorptance,

λ = wavelength, and λ_1 and λ_2 , define the wavelength limits of the source.

10.6.6 It is clearly evident that to obtain meaningful data and to preclude extreme errors during a steady state a/ε measurement using solar simulation, it is necessary to use either a filtered xenon compact arc lamp or a carbon arc source.

11. Keywords

11.1 calorimetry; emittance; infrared emittance; material radiative property; radiative heat transfer; solar absorptance; spacecraft thermal control; spectral normal emittance; thermal radiation

TABLE 1 Typical Absorptance Values for Various Solar Simulation Sources

Coating	Xenon Compact Arc Un-filtered	Xenon Compact Arc Filtered	Mercury-Xenon 2.5 kW Compact Short Arc	Krypton Compact Arc	Carbon Arc	Low Pressure Mercury (500 W)	Solar Irradiance
3M Velvet Black Paint, 100 Series	0.972	0.972	0.972	0.973	0.972	0.971	0.972
ZnO pigment in methylsilicone binder over GE primer on buffed Al (S-13)	0.165	0.188	0.345	0.226	0.181	0.502	0.189
TiO ₂ pigment in methylsilicone binder over cat-a-lac white primer on buffed Al (Dow Corning Q92-090)	0.140	0.160	0.303	0.193	0.149	0.465	0.163
Evaporated gold	0.145	0.186	0.235	0.128	0.166	0.385	0.198
Evaporated silver	0.041	0.040	0.134	0.060	0.032	0.285	0.050
Evaporated aluminum	0.080	0.074	0.065	0.073	0.071	0.073	0.076

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